

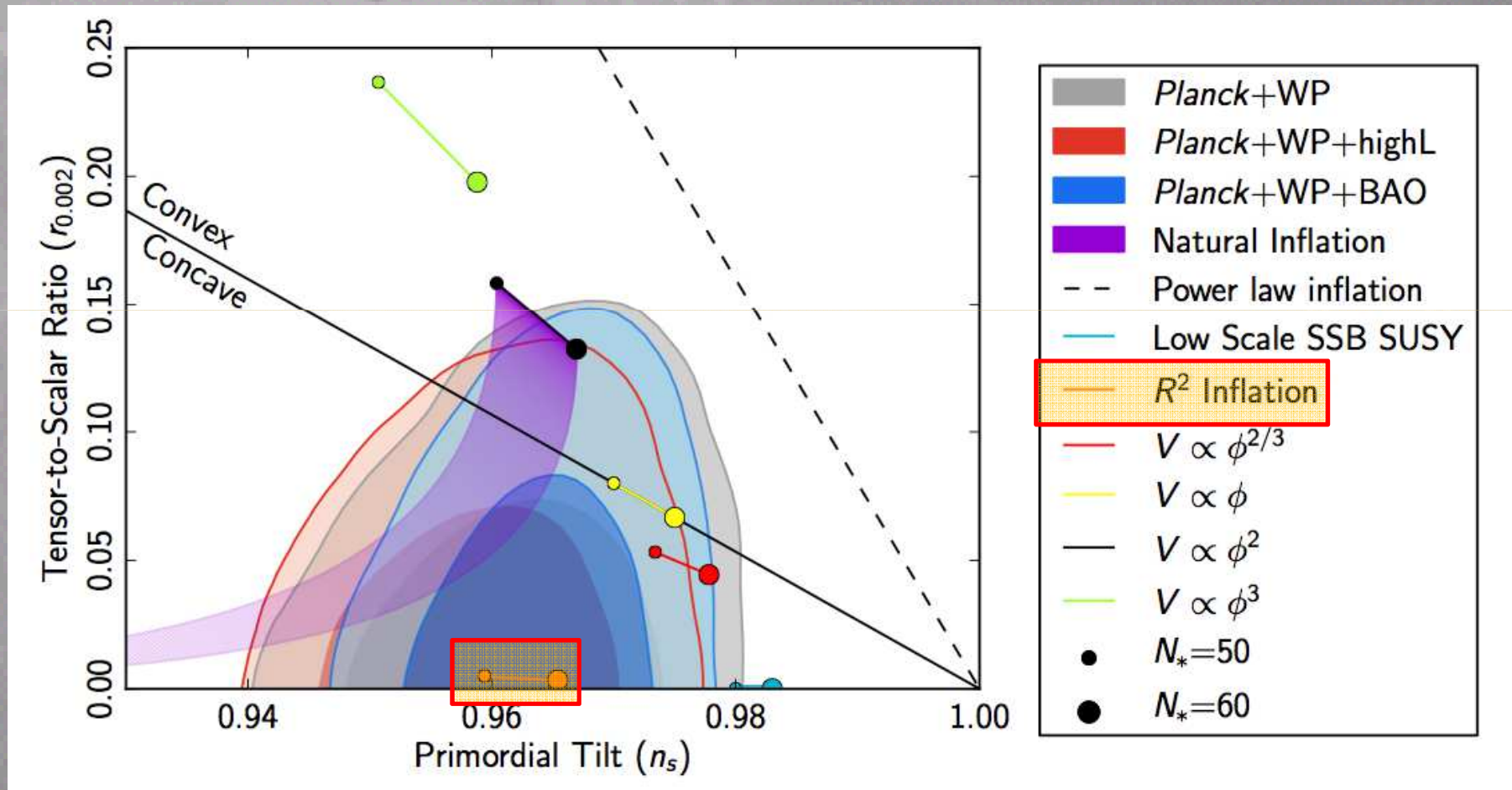
$R + R^2$ supergravity and off-shell multiplets

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mostly based on collaborations
with **R. Kallosh** and **S. Ferrara**

Dubna workshop
‘What next?’
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Planck 2013 results; constraints on inflation



Planck collaboration, 1303.5082

Starobinsky inflation and higher derivative terms in gravity

- Higher derivative terms to replace initial singularity (Starobinsky, 1980);
rewritten as $R+R^2$: Kofman, Linde, Starobinsky, 1985

- Correcting Einstein theory (2nd order Lagrangian) with 4th order terms

$$S = \int d^4x \sqrt{g} L, \quad L = \frac{1}{2\kappa^2} R$$
$$\kappa^2 = 8\pi G$$

$$L = \frac{1}{2\kappa^2} R + a R_{\mu\nu\rho\sigma} R^{\mu\nu\rho\sigma} + b R_{\mu\nu} R^{\mu\nu} + c R^2$$

- These are also first string corrections (and also studied for black hole entropy, corrections to AdS/CFT, quantum loops, ...)

Plan

1. Gravity: 4th order derivative terms and conformal structure
2. N=1 supergravity: auxiliary fields and supersymmetrization of $R+R^2$ terms
3. N=2 auxiliary fields and expectations for dual theories to $R+R^2$ supergravity
4. N=4 superconformal structure
5. Conclusions

4th order terms in gravity

- Euler density $L = R_{\mu\nu\rho\sigma}R^{\mu\nu\rho\sigma} - 4R_{\mu\nu}R^{\mu\nu} + R^2$
is a topological invariant
→ does not contribute to field equations.

- Conformal invariant
$$\begin{aligned} L &= C_{\mu\nu\rho\sigma}C^{\mu\nu\rho\sigma} \\ &= R_{\mu\nu\rho\sigma}R^{\mu\nu\rho\sigma} - 2R_{\mu\nu}R^{\mu\nu} + \frac{1}{3}R^2 \\ &\simeq 2(R_{\mu\nu}R^{\mu\nu} - \frac{1}{3}R^2) \end{aligned}$$

- We parametrize
$$\begin{aligned} L &= \frac{1}{2\kappa^2}R + a R_{\mu\nu\rho\sigma}R^{\mu\nu\rho\sigma} + b R_{\mu\nu}R^{\mu\nu} + c R^2 \\ &= \frac{1}{2\kappa^2}R + \alpha R^2 + \beta(R_{\mu\nu}R^{\mu\nu} - \frac{1}{3}R^2) \end{aligned}$$

- K. Stelle, 1977:

- without α and β : **massless spin 2**
- with α : additional **physical spin 0** with $m_0^{-2} = 12 \alpha \kappa^2$
- with β : additional **ghost spin 2** with $m_2^{-2} = -2\beta \kappa^2$

Degrees of freedom (dof)

- **On-shell** dof = # helicity states (4 dim)
 - massless: **2** for every spin except $s=0$: **1** state
 - massive : **$2s+1$**
- **Off-shell** dof = # field components – # gauge transform
 - real scalar: **1**
 - Majorana fermion: **4**
 - Gauge vector: $4 - 1 = **3**$
 - metric field : $16 - 6 - 4 = **6**$ (local Lorentz and translations)
 - gravitino : $16 - 4 = **12**$ (local susy)
- **Supersymmetry**: as well on-shell as off-shell
bosonic dof = # fermionic dof
(only known for $N=1,2$):
necessity of auxiliary fields (0 dof on shell)

Particle states

$$L = \frac{1}{2\kappa^2}R + \alpha R^2 + \beta(R_{\mu\nu}R^{\mu\nu} - \frac{1}{3}R^2)$$

massless
spin 2

massive
spin 0

conformal →
ghost massive spin 2

- A massless spin 2 has 2 degrees of freedom (dof)
- Off-shell metric field has 6 dof
- Conformal metric field (Weyl field) has 5 dof
- A massive spin 2 has 5 dof

■ **Poincaré action** from
Weyl
+ **compensating scalar** :

In higher derivative action
ghost
physical scalar

Goal of this talk

- Show how this structure is valid with supersymmetry
- auxiliary fields (off shell dof) become propagating
⇒ structure of auxiliary fields already determines content of massive sector
⇒ different auxiliary fields → different massive sector
- In conformal structure: different compensators
 - in $N=1$: understanding of two inequivalent generalizations
 - in $N=2$: there are 3 different generalizations with different compensating multiplets, and we thus expect different massive sectors for higher-derivative theories

Conformal structure: Weyl multiplet

| | | |
|----------|-------------------------------------|---|
| P_a | 4 translations | e_μ^a |
| M_{ab} | 6 Lorentz rotations | ω_μ^{ab} |
| D | 1 dilatation | b_μ |
| K_a | 4 special conformal transformations | f_μ^a |

Constraints determine two gauge fields

$$\omega_\mu^{ab} = \omega_\mu^{ab}(e, b)$$

$$f_\mu^a = -\frac{1}{4}R_\mu^a + \frac{1}{24}e_\mu^a R$$

‘Weyl multiplet’: e_μ^a, b_μ

$$\text{dof: } e_\mu^a : 16 - 4 - 6 - 1 = 5$$

$$b_\mu : 4 - 4 = 0$$

$$\delta_K b_\mu = e_{\mu a} \lambda_K^a$$

gauge choice : $b_\mu = 0$

Conformal action

Consider a scalar. We have to define dilatation transformation

$$\delta_D \phi = w \lambda_D \phi$$

w parameter: 'Weyl weight'

$$\square^C \phi \equiv \eta^{ab} \mathcal{D}_b \mathcal{D}_a \phi = \square \phi + 2w e^{a\mu} f_{\mu a} \phi \quad e^{\mu a} f_{\mu a} = -\frac{1}{12} R$$

$$w = 1 \quad \square^C \phi = \square \phi - \frac{1}{6} R \phi \quad \delta_{D, \kappa} \square^C \phi = 3 \lambda_D \square^C \phi.$$

$$I = -\frac{1}{2} \int d^4 x e \phi \square^C \phi = \int d^4 x e \left(\frac{1}{2} g^{\mu\nu} (\partial_\mu \phi) (\partial_\nu \phi) + \frac{1}{12} R \phi^2 \right)$$



wrong sign kinetic terms;
but this is not physical

gauge fix dilatations $\phi^2 = \frac{6}{\kappa^2} : I = \int d^4 x e \frac{1}{2\kappa^2} R$

Higher order action

- The non-conformal action $L = \frac{1}{2\kappa^2}R + \alpha R^2$

with previous gauge $\phi = \frac{\sqrt{6}}{\kappa}$, $\square^C \phi = -\frac{1}{\sqrt{6}\kappa}R$, $\frac{\square^C \phi}{\phi} = -\frac{1}{6}R$

is the conformal action

$$L = -\frac{1}{2}\phi \square^C \phi + 36\alpha \left(\frac{\square^C \phi}{\phi}\right)^2$$

- Conformal dualization:

$$L = -\frac{1}{2}\phi \square^C \phi + \sigma \left(\chi - \frac{\square^C \phi}{\phi}\right) + 36\alpha \chi^2$$

eliminate χ :

$$L = -\left(\frac{1}{2}\phi^2 + \sigma\right) \frac{\square^C \phi}{\phi} - \frac{1}{144\alpha} \sigma^2$$

D - gauge:

$$\frac{1}{2}\phi^2 + \sigma = \frac{3}{\kappa^2}$$

$$L = -\frac{3}{\kappa^2} \left(\frac{\square^C \phi}{\phi} - \frac{1}{6}R\right) - \frac{1}{144\alpha} \left(\frac{3}{\kappa^2} - \frac{1}{2}\phi^2\right)^2$$

$$\simeq \frac{1}{2\kappa^2}R - \frac{3(\partial_\mu \phi)(\partial^\mu \phi)}{(\kappa\phi)^2} - \frac{1}{16\alpha\kappa^4} \left(1 - \frac{1}{6}\kappa^2\phi^2\right)^2$$

physical normalization

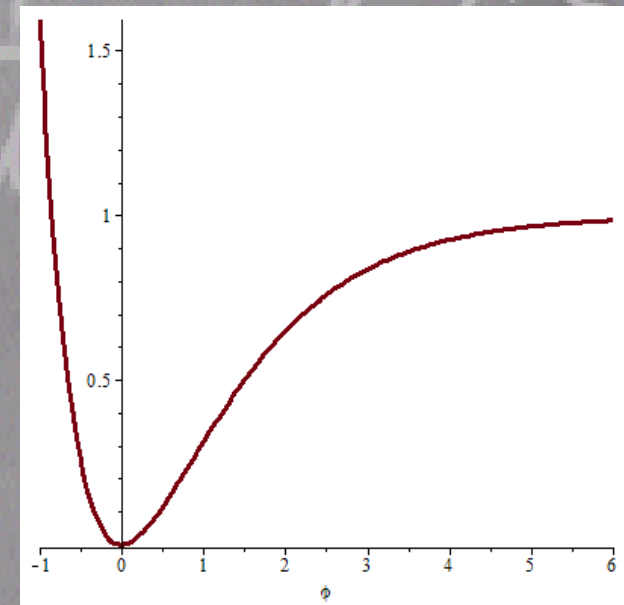
$$\kappa\phi = \sqrt{6} \exp\left(-\frac{1}{\sqrt{6}}\kappa\varphi\right)$$

$$L = \frac{1}{2\kappa^2}R - \frac{1}{2}(\partial_\mu \varphi)(\partial^\mu \varphi) - \frac{1}{16\alpha\kappa^4} \left[1 - \exp\left(-\sqrt{\frac{2}{3}}\kappa\varphi\right)\right]^2$$

Starobinsky model result

$$L = \frac{1}{2\kappa^2}R - \frac{1}{2}(\partial_\mu\varphi)(\partial^\mu\varphi) - \frac{1}{16\alpha\kappa^4} \left[1 - \exp\left(-\sqrt{\frac{2}{3}}\kappa\varphi\right) \right]^2$$

- Thus: $R + R^2$ without conformal term: equivalent to one **physical propagating massive compensating field**.
- Original result: **B. Whitt, 1984**
- Scalar with **positive kinetic terms** ('scalaron'), and mass as in seminal work of K. Stelle, potential suitable for slow roll inflation to Minkowski vacuum



Schematic

■ off-shell Poincaré (6 dof) =

conformal spin 2 (5 dof) + compensating scalar (1 dof)

↓
Weyl action

↓
dual of R^2 action

massive spin 2 ghost + massive spin 0 physical

■ on-shell Poincaré (2 dof)

R action

N=1 supergravity

■ Massless multiplets:

■ Massive multiplets

| spin | # | dof/particle | bosonic | fermionic |
|--------------|---|--------------|---------|-----------|
| 2 | 1 | 2 | 2 | |
| 3/2 | 1 | 2 | | 2 |
| <i>TOTAL</i> | | | 2 | 2 |

| spin | # | dof/particle | bosonic | fermionic |
|--------------|---|--------------|---------|-----------|
| 2 | 1 | 5 | 5 | |
| 3/2 | 2 | 4 | | 8 |
| 1 | 1 | 3 | 3 | |
| <i>TOTAL</i> | | | 8 | 8 |

| spin | # | dof/particle | bosonic | fermionic |
|--------------|---|--------------|---------|-----------|
| 1 | 1 | 2 | 2 | |
| 1/2 | 1 | 2 | | 2 |
| <i>TOTAL</i> | | | 2 | 2 |

| spin | # | dof/particle | bosonic | fermionic |
|--------------|---|--------------|---------|-----------|
| 1 | 1 | 3 | 3 | |
| 1/2 | 2 | 2 | | 4 |
| 0 | 1 | 1 | 1 | |
| <i>TOTAL</i> | | | 4 | 4 |

| spin | # | dof/particle | bosonic | fermionic |
|--------------|---|--------------|---------|-----------|
| 1/2 | 1 | 2 | | 2 |
| 0 | 2 | 1 | 2 | |
| <i>TOTAL</i> | | | 2 | 2 |

| spin | # | dof/particle | bosonic | fermionic |
|--------------|---|--------------|---------|-----------|
| 1/2 | 1 | 2 | | 2 |
| 0 | 2 | 1 | 2 | |
| <i>TOTAL</i> | | | 2 | 2 |

N=1 D=4 Superconformal gauge fields and the Weyl multiplet

| | | | | | | |
|-----------|-------------------|---------|-----------|---------|------------|------------|
| P_a | M_{ab} | D | K_a | U(1) | Q | S |
| 4 | 6 | 1 | 4 | 1 | 4 | 4 |
| e_μ^a | ω_μ^{ab} | b_μ | f_μ^a | A_μ | ψ_μ | ϕ_μ |

determined by constraints

‘Weyl multiplet’: $e_\mu^a, b_\mu, A_\mu, \psi_\mu$

| field | # components | spin 2 | spin $\frac{3}{2}$ | spin 1 |
|------------|----------------------|--------|--------------------|--------|
| e_μ^a | $16 - 4 - 6 - 1 = 5$ | 1 | | |
| b_μ | $4 - 4 = 0$ | | | |
| A_μ | $4 - 1 = 3$ | | | 1 |
| ψ_μ | $16 - 4 - 4 = 8$ | | 2 | |

is the massive mult.

| spin | # | dof/particle |
|------|---|--------------|
| 2 | 1 | 5 |
| 3/2 | 2 | 4 |
| 1 | 1 | 3 |

explicit in

Supergravity



Daniel Freedman and Antoine Van Proeyen

CAMBRIDGE

Superconformal methods for cosmology, see review R. Kallosh, 1402.0527

To super-Poincaré

- super-Poincaré physical fields:

$$e_{\mu}^a : 16 - 4 - 6 = 6; \quad \psi_{\mu} : 16 - 4 = 12$$

- *We need auxiliary fields.*

- Other way: compare with conformal

$$e_{\mu}^a : 16 - 4 - 6 - \overset{D}{1} = 5; \quad \psi_{\mu} : 16 - 4 - \overset{S}{4} = 8$$

$$A_{\mu} : 4 - \underset{U(1)}{1} = 5$$

We need compensators: the minimal (‘old minimal’)

- chiral mult. $(Z, \Omega, F) \rightarrow$ aux. fields $A_{\mu}(4)$ en $F(2)$
- linear mult. $(L, \varphi, E_{\mu\nu}) \rightarrow$ aux. fields $A_{\mu}(3)$ en $E_{\mu\nu}(3)$

‘new minimal’

In higher derivative theories

- If Weyl action:

Weyl multiplet is massive spin 2 ghost multiplet.

- If R^2 terms: after dualization: compensating multiplet becomes physical massive multiplet

- old minimal (off-shell chiral mult);

| field | # components | spin $\frac{1}{2}$ | spin 0 |
|----------|--------------|--------------------|--------|
| X | 2 | | 2 |
| F | 2 | | 2 |
| Ω | 4 | 2 | |

compare massive chiral mult.

| spin | # |
|-------|---|
| $1/2$ | 1 |
| 0 | 2 |

With old minimal we get 2 massive chiral multiplets !

Superconformal tensor calculus for old-minimal supergravity

- Poincaré supergravity: with compensating multiplet $-[S_0\bar{S}_0]_D$
- R^2 action results from kinetic multiplet $\mathcal{R}(S_0)$ that contains the curvature:

$$[-S_0\bar{S}_0 + \mathcal{R}(S_0)\overline{\mathcal{R}(S_0)}]_D$$

- Dualization

$$\begin{aligned} & [-S_0\bar{S}_0 + S\bar{S}]_D + [S_0^2\sigma(S - \mathcal{R}(S_0))]_F \\ &= [-S_0\bar{S}_0(1 + \sigma + \bar{\sigma}) + S\bar{S}]_D + [S_0^2\sigma S]_F \end{aligned}$$

S_0 : compensating: S and σ physical + potential

Conclusion on old minimal set

- Dual of $R+R^2$ supergravity is a matter coupling with 2 chiral multiplets

(S. Cecotti; 1987; R. Kallosh, S. Ferrara, AVP, 1309.4052)

- Requires stabilization of other scalars

(J. Ellis, D. Nanopoulos, D. Olive, 1305.1247; 1307.3537;
R. Kallosh and A. Linde, 1306.3214)

New minimal set of auxiliary fields

- without higher derivatives: dual to old minimal:
S. Ferrara, L. Girardello, T. Kugo and AVP, 1983

| field | # components | spin 1 | spin $\frac{1}{2}$ | spin 0 |
|-----------|--------------|--------|--------------------|--------|
| L | 1 | | | 1 |
| E_a | 3 | 1 | | |
| φ | 4 | | 2 | |

| spin | # |
|------|---|
| 1 | 1 |
| 1/2 | 2 |
| 0 | 1 |

- R^2 term gives different result: massive vector multiplet:
S. Cecotti, S. Ferrara, M. Porrati and S. Sabharwal, 1987;
S. Ferrara, R. Kallosh, A. Linde and M. Porrati, 1307.7696.
- This massive multiplet was constructed in AVP, 1980.
- Only 1 scalar, does not need stabilization.
- Further possibilities for inflaton field, Kähler structure and gauged isometries in
S. Ferrara, P. Frè, A. Sorin, 1311.5059 en 1401.1201

Conclusions

- Conformal construction of supergravity shows the content of higher derivative actions.
- Different compensating multiplets imply different massive multiplets that are dual to $R+R^2$ actions.
- $N=1$: known since old papers of S. Cecotti and S. Ferrara, M. Porrati and S. Sabharwal, 1987. Especially new-minimal interesting dual: a massive vector multiplet with only one physical scalar. Recently studied in more detail in view of cosmology application.
- $N=2$: different auxiliary field formulations known. They may lead to different physical $R+R^2$ theories.
- $N=4$: probably no $R+R^2$ theory.